



2x2 Chicago Meeting

April 2, 2024 Angela White and Elise Hinkle



2x2 Paper Update

Link to Overleaf Draft:

https://www.overleaf.com/6419191385gzxrjygksbcb#1adc5e

- Introduction, Hardware, and Commissioning are drafted
- Currently communicating with cosmic + rock muon simulation folks for Backgrounds section
- Updated beam flux plots currently in production
- Need better event displays-Elise has begun pushing this in working groups

Any and all comments welcome in the draft above!

Other Event Display

- Made stop-gap event display for placeholder in paper draft
- Used it in presentation in analysis meeting to advocate for more event displays \rightarrow hopefully, there will be more/better event displays coming



2x2 LRS Warm Commissioning

Does the Light Output Make Sense?



Observe: No light signal at tick 200–just dark counts throughout

Test 1: Over 1005 Events, Avg. Maximum by Channel



Observe:

Only concerned with global maximum

Single p.e. dark count: ~ 300 ADC counts

Test 1: Over 1005 Events, Avg. Maximum by Channel



Observe: Max. dark count is ~ 1 p.e. In Mod-0, and ~ 3 p.e. in Mods 1, 2, &3

Test 2: Over 1005 Events, Avg. Amplitude by Channel



Observe:

Only concerned with global average

Average this value for each of 384 SiPM channels over 1005 events...

Test 2: Over 1005 Events, Avg. Amplitude by Channel



Observe: Not very different. ACLs, in general, are noisier. Mod 3 ACLs are non-responsive.

Test 3: Over 1005 Events, Avg. Ratio of Noise Amp Across Wvfm



Observe:

Take ratio: Avg [800:1000] / Avg [0:200]

Test 2: Over 1005 Events, Avg. Amplitude by Channel



Observe: On average, there are as many dark counts early in the waveform as there are late

Test 4: Over 1005 Events, Avg. Dark Count Rate



Observe:

Most dark counts appear to be single p.e.

Test 2: Over 1005 Events, Avg. Amplitude by Channel



Observe: On average, 30 dark counts / 16 µs (1.9 MHz), lower for Mod-0

Light Noise FFTs:

Main Takeaway: Dark Counts in warm drown out electronics noise

• See: 10 MHz peak



Light Noise FFTs:

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Updates on ML Reco Benchmarking

- Gave additional presentation at ND Prototypes Analysis Meeting March 21, 2024
- Some updates from previously shown studies + additional suggestions

Current Work – Full Reco Benchmarking

- Still want to look at calibration-file-level comparisons of data/MC using Bern Module data and simulation (see my <u>presentation last week</u> for update on Bern module cosmics simulation status)
- Also want to investigate full proton reconstruction using CAFs by comparing reco and true particle kinematics (see <u>other presentation last week</u> for first results)
- As CAFs currently only contain **ML Reco** information, this is the reconstruction I'm evaluating



Sample Details

- Used MiniRun4.5 Beta 3 CAFs (200 files)
- As ML Reco has some known PID issues, I look at all **reconstructed** charged track-like particles and also just **reconstructed** protons in comparison to **best match true particles**
- Cut on ML Reco "Overlap" variable such that require reco/true match to have >=0.5 overlap
- No throughgoing tracks
- No tracks with reco start or end points at the upstream edge of detector (within 1.0 cm)
- •Overall, 19594 charged tracks and 5726 protons



Charged Track and Proton Length



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DEEP UNDERGROUND NEUTRINO EXPERIMENT

• More short true tracks



 Many more than expected reco/true matches with >10 cm track length differences

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Charged Track and Proton Length Comparison



 True match tracks more likely to be shorter than reconstructed tracks vs. longer

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Charged Track Start Position



True vs. ML Reco Track Start Position for Reconstructed Charged Track Sample

- Significant differences in x-coordinate distribution
- Large spikes at edges for ML Reco in y, z



Proton Start Position



True vs. ML Reco Track Start Position for Reconstructed Proton Sample

- Significant differences in x-coordinate distribution
- Large spikes at edges for ML Reco in y, z



Charged Track End Position



True vs. ML Reco Track End Position for Reconstructed Charged Track Sample

• Significant differences in x-coordinate distribution



Proton End Position



True vs. ML Reco Track End Position for Reconstructed Proton Sample

• Significant differences in x-coordinate distribution





(Absolute value of the cosine of the)

Charged Track Angle w.r.t Beam



- Binned by reconstructed track length in 10 cm bins
- •Note: log scale on y-axis
- For shorter tracks, clear difference in true vs. reco distributions





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NEUTRINO EXPERIMENT







- Binned by reconstructed track length in 10 cm bins
- Note: normal scale on y-axis
- For all tracks, clear difference in true vs. reco distributions

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Difference in Abs. Cos. of Pixel Plane Angle



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•Line at +0.75 difference
in filtered here and looked at true vs. reco tracks above this difference threshold

Ex: Charged Track #1

 Best alignment on "pixel plane" view (bottom left)



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NEUTRINO EXPERIMENT

"Full Event" #1

- Reconstructed tracks shown:
 - All reco tracks from same file, spill, and reco interaction as reco track in last slide
- True tracks shown:
 - All reco tracks from same file, spill, and true interaction as true track in last slide



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NEUTRINO EXPERIMENT

Ex: Charged Track #2

 Best alignment on "pixel plane" view (bottom left)



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NEUTRINO EXPERIMENT

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"Full Event" #2

Reconstructed tracks shown:

- All reco tracks from same file, spill, and reco interaction as reco track in last slide
- True tracks shown:
 - All reco tracks from same file, spill, and true interaction as true track in last slide



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Ex: Charged Track #3

- Best alignment on "pixel plane" view (bottom left)
- Note that true match end and ML Reco start match better than true/reco start or true/reco end



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NEUTRINO EXPERIMENT

"Full Event" #3

Reconstructed tracks shown:

 All reco tracks from same file, spill, and reco interaction as reco track in last slide

• True tracks shown:

• All reco tracks from same file, spill, and true interaction as true track in last slide



NEUTRINO EXPERIMENT

Ex: Charged Track #4

 Best alignment on "pixel plane" view (bottom left)



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DEEP UNDERGROUND

NEUTRINO EXPERIMENT

"Full Event" #4

• **Reconstructed** tracks shown:

- All reco tracks from same file, spill, and reco interaction as reco track in last slide
- True tracks shown:
 - All reco tracks from same file, spill, and true interaction as true track in last slide



Track Multiplicity at Vertex



• First bin may be cases where true particle match is shower-like

• In the future, will look at kinematics by true track multiplicity at vertex to get a better understanding of reconstruction fidelity in high activity environments

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NEUTRINO EXPERIMENT

Future Studies

- Break down plots I showed in terms of different variables (e.g. by charged track multiplicity at the vertex, by start/end position, etc.) to identify specific failure modes
- Similar studies w/ reflowed Bern data/new cosmics samples run through ML Reco
- Look at proton thresholds using a sample of true protons
- Create **samples of TRUE protons and charged tracks** and make plots similar to what I've shown here
- Make efficiency vs. purity plots with reco protons, charged tracks
- Make plots such that they can be **easily reproduced** for new iterations of ML Reco (i.e. when it is retrained)
- Make more informative full event display to see all activity vs. single set of matched tracks

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Additional Comments from Others

- Look at angles based on Cartesian coordinates
- Look at differences true vs. reco angles and start/end positions
- Look at events for particular failure modes in official ML Reco event display
- Look at events with different "overlap" amounts true vs. reco